

Generalization of Z string and its application

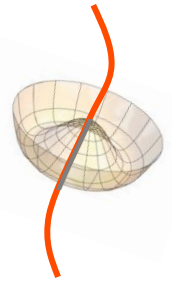
Yukihiro Kanda (Nagoya University)

Based on

YK, N. Maekawa, PRD 107 (2023) 9, 096007 [arXiv: 2303.09517]

Gravitational Wave Probes of Physics Beyond Standard Model
2023/11/9, Osaka Metropolitan University

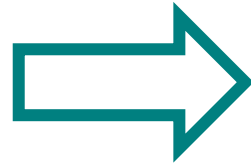
Introduction



Cosmic strings are

- linear defects produced during some symmetry breakings
- probed by gravitational wave observations
- cannot be produced in the SM

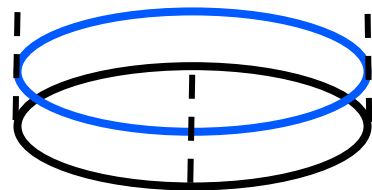
Find cosmic string
through GWs



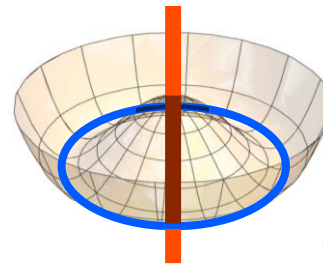
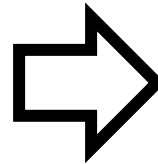
Physics BSM
is detected

Cosmic strings are produced when the vacuum manifold of symmetry breaking has a **non-contractible loop**. [Kibble (1976)]

e.g. $U(1) \rightarrow \times$



Vacuum manifold $\mathcal{V} \simeq S^1$

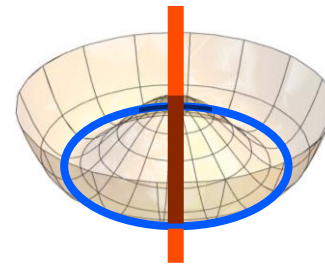
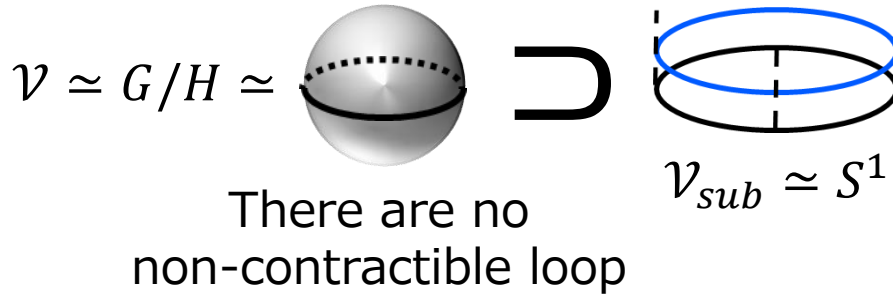


Classical solution
having a **linear excited region**

Embedded string

Can cosmic strings not be produced if there are no non-contractible loop on the vacuum manifold? ➔ NO!

When $G \rightarrow H$,

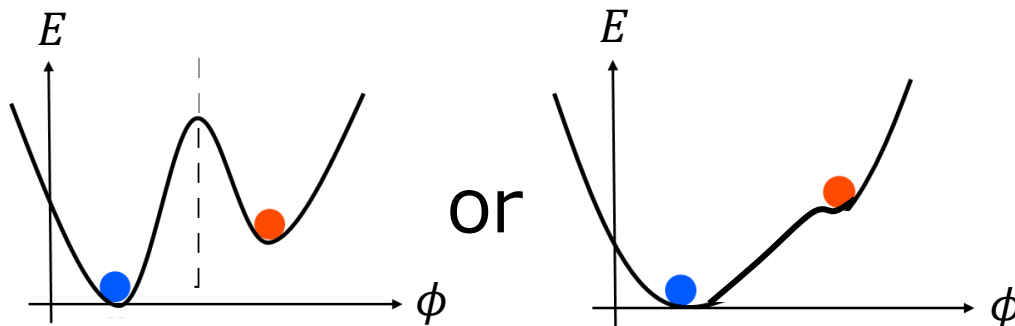


String like classical solution

“Embedded string”

[Vachaspati, Barriola (1992)]
[Vachaspati, Barriola, Bucher (1994)]

Embedded string's stability is not topologically guaranteed



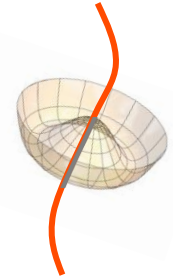
or

- : constant vacuum solution
- : Embedded string solution



Determined by the parameters

Motivation



Embedded string is classically stable.

⇒ it is produced during a symmetry breaking.

However, it has been studied only for one embedded string (Z string).

[James, Perivolaropoulos, Vachaspati (1993)]

To probe physics BSM through GW from cosmic strings, it is important to study the stability of other embedded strings.

Today's talk

- We have studied generalization of the Z string, and found that its stability is determined only by the mass ratios of the Higgs and the gauge bosons.
- We have applied this result to cases of gauge group unification.
YK, N. Maekawa (2023)
- We comments on the relevance to the recent result of pulsar timing array collaborations.
Our ongoing work

Z string

$$SU(2) \times U(1) \xrightarrow{H: (\mathbf{2}, 1/2)} U(1)$$

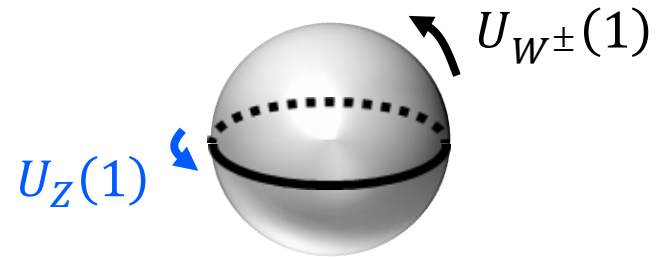
$\mathcal{V} \simeq SU(2) \times U(1)/U(1) \simeq S^3$
No non-contractible loops on S^3

$$U \quad U$$

$$U_Z(1) \longrightarrow \times$$

related to the Z boson

Consider embedded string



Z-string [Vachaspati (1992)]

$$H = \begin{pmatrix} 0 \\ f(r)ve^{i\theta} \end{pmatrix}, \quad \vec{Z} = -\frac{2z(r)}{\alpha r} \vec{e}_\theta, \quad (\text{others}) = 0$$

in cylindrical coordinate (r, θ, z) , $\alpha^2 = g_1^2 + g_2^2$

- $f(r)ve^{i\theta}, -\frac{2z(r)}{\alpha r}$ are string solution of $U(1)$ breaking

$$(f(0) = z(0) = 0, f(\infty) = z(\infty) = 1)$$

Stability of the Z string

[James, Perivolaropoulos, Vachaspati (1993)]

Consider all the perturbations from the Z-string

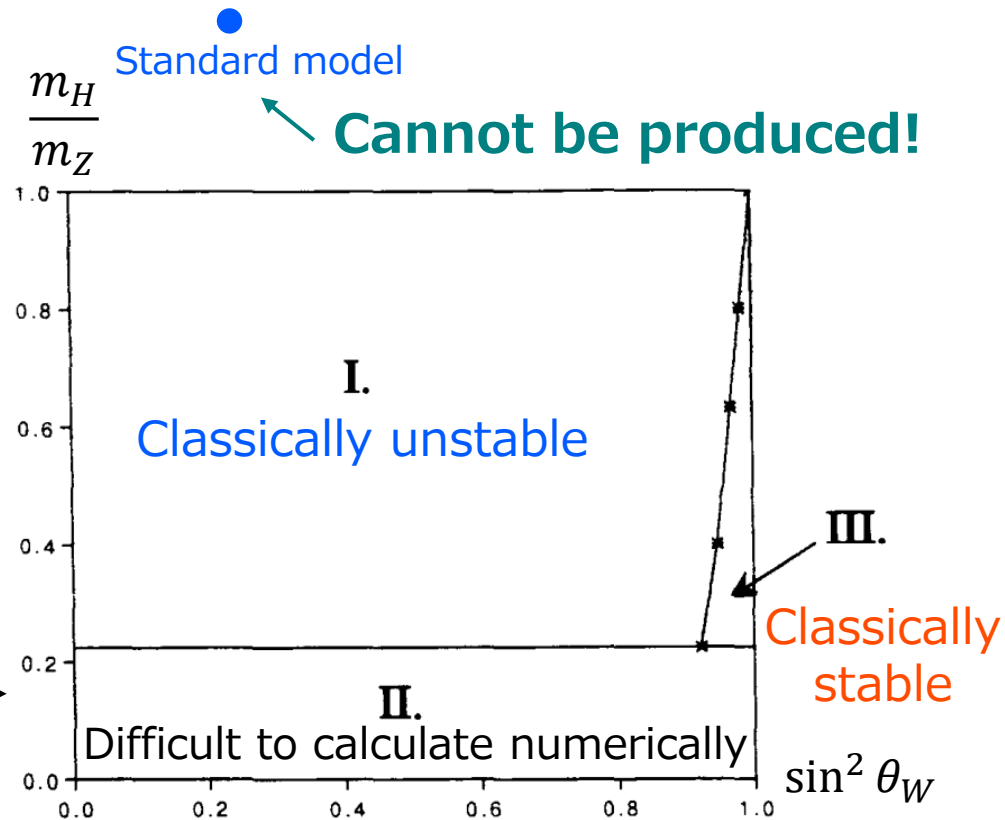
$$H = \begin{pmatrix} \phi(x) \\ f(r)ve^{i\theta} + \delta h(x) \end{pmatrix}, \vec{Z} = -\frac{z(r)}{\alpha r} \vec{e}_\theta + \delta \vec{Z}(x), \vec{W}^\pm(x), \vec{A}(x)$$

↓ Calculate the variations of the energy $\delta\mu$ and find modes decreasing it

Only one mode can destabilize the Z string

$$\delta\mu = \int R dR \zeta(R) \mathcal{O} \left(R; \frac{m_H}{m_Z}, \theta_W \right) \zeta(R)$$

$R \equiv \frac{\alpha v}{2} r, \tan \theta_W \equiv g_1/g_2$
 m_H, m_Z : mass of scalar and Z boson
 ζ : perturbation mode



Overview of our work

- The stability of the Z string was studied in 1990s.
- The stability of other embedded strings were not studied.
(∵ motivated by the electroweak breaking)



But now,

- Models beyond the standard model have various symmetry breakings.
- If cosmic strings are produced during a symmetry breaking, they can be probed by GW observation.

Our work

We consider the generalization of the Z string for $SU(N) \times U(1)_X \rightarrow SU(N-1) \times U(1)_Q$ and study the stability of them.

Generalization of Z string

We consider $SU(N) \times U(1)_X \xrightarrow{\phi: \left(N, \frac{1}{2}\right)} SU(N-1) \times U(1)_Q$

Scalar potential: $V(\phi) = \lambda(|\phi|^2 - v^2)^2 \longleftarrow v \simeq S^{2N-1}$

No non-contractible loop

➔ There is a neutral massive gauge boson \tilde{Z}_μ

$$\tilde{Z}_\mu \equiv \sqrt{\frac{2(N-1)}{N}} \frac{g_N}{\alpha_N} G_\mu^{N^2-1} - \frac{g_1}{\alpha_N} B_\mu$$



Make an embedded string

$$\left(\begin{array}{l} G_\mu^a, B_\mu: SU(N), U(1) \text{ gauge bosons} \\ T^{N^2-1} = \frac{1}{\sqrt{2N(N-1)}} \text{diag}(1, \dots, 1, 1-N) \\ \alpha_N^2 \equiv \frac{2(N-1)}{N} g_N^2 + g_1^2 \end{array} \right)$$

Generalized Z-string

$$\phi = \begin{pmatrix} 0 \\ \vdots \\ 0 \\ f(r) v e^{i\theta} \end{pmatrix}$$

$$\vec{\tilde{Z}} = -\frac{2z(r)}{\alpha_N r} \vec{e}_\theta, \quad (\text{others})=0$$

$$f(0) = z(0) = 0, f(\infty) = z(\infty) = 1$$

Note that it is the Z-string when $N = 2$

Calculation of the stability

Scalar: $\phi(x) = \begin{pmatrix} \phi_1(x) \\ \vdots \\ \phi_{N-1}(x) \\ f(r)e^{in\theta} + \delta\phi(x) \end{pmatrix}$, Gauge boson: $\begin{pmatrix} \vec{G}^a(x) & \vec{G}^+(x) \\ \vec{G}^-(x) & \end{pmatrix}$,

 : $SU(N-1)$ adjoint
 : $SU(N-1)$ fundamental
 : $SU(N-1)$ singlet

$$\vec{Z} = -\frac{z(r)}{\alpha_N r} \vec{e}_\theta + \delta\vec{Z}(x),$$

$$\vec{A}(x) \equiv \frac{g_1}{\alpha_N} \vec{G}^{N^2-1} + \sqrt{\frac{2(N-1)}{N}} \frac{g_N}{\alpha_N} \vec{B}$$

Diagonal part

↓ Calculate the variations of the energy $\delta\mu$

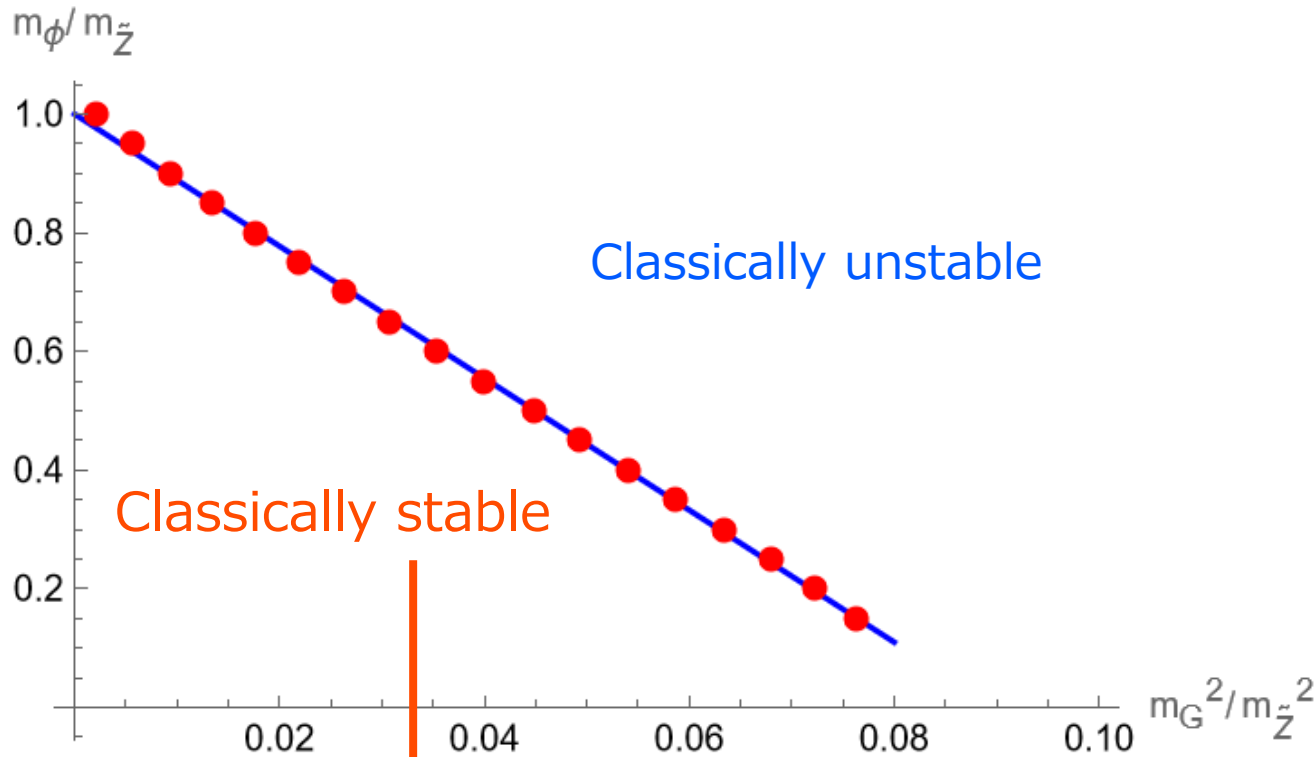
Only fundamental modes can destabilize the generalized Z string.
 $\delta\mu$ is divided into $N-1$ parts which are similar to $\delta\mu$ of the Z string.

$$\delta\mu = \sum_{k=1}^{N-1} \delta\mu_k(r; m_\phi/m_{\vec{Z}}, m_G/m_{\vec{Z}})$$

$m_\phi, m_{\vec{Z}}, m_G$: the mass of scalar, neutral gauge boson, charged gauge boson
 θ_G : mixing angle

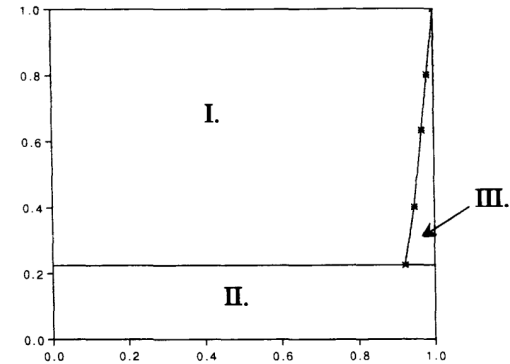
Consistent with the Z string (cf. $\frac{m_G}{m_{\vec{Z}}} = \sqrt{\frac{N}{2(N-1)}} \cos \theta_G$)

Result

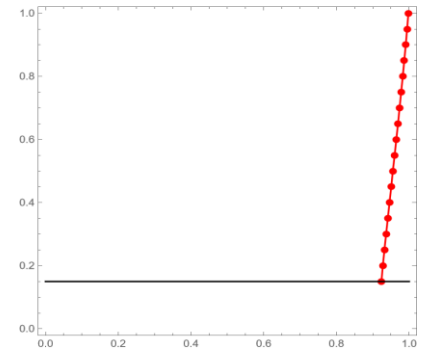


$N = 2$

[James, Perivolaropoulos, Vachaspati (1993)]



Ours result



approximately
evaluation

$$\frac{m_\phi}{m_{\tilde{Z}}} \leq 1 - 11 \frac{m_G^2}{m_{\tilde{Z}}^2} \Leftrightarrow g_1 \geq \sqrt{\frac{11}{1 - m_\phi/m_{\tilde{Z}}} - \frac{2(N-1)}{N}} g_N$$

Application for unification

We consider the case that $SU(N)$ and $U(1)$ have the same origin

$$\begin{array}{ccc}
 \phi = (N, q, \mathbf{1}) \Big|_{g_1'} & = & (N, 1/2, \mathbf{1}) \Big|_{g_1} \\
 \downarrow & & \\
 G \rightarrow \cdots \rightarrow SU(N) \times U(1) \times H & \rightarrow & SU(N-1) \times U(1) \times H \\
 \\
 g_U = g_N = g_1' & \xrightarrow{\text{RG running}} & g_N = \alpha_{RG} g_1' = \frac{\alpha_{RG}}{2q} g_1
 \end{array}$$

The generalized Z-strings are formed when g_N and g_1 satisfy

$$g_1 \geq \sqrt{\frac{11}{1 - m_\phi/m_{\tilde{z}}} - \frac{2(N-1)}{N}} g_N \quad \Rightarrow \quad \boxed{q^2 \geq \alpha_{RG}^2 \left[\frac{2.75}{1 - m_\phi/m_{\tilde{z}}} - \frac{N-1}{2N} \right]}$$

Condition for the rep. of ϕ in G

We apply it for

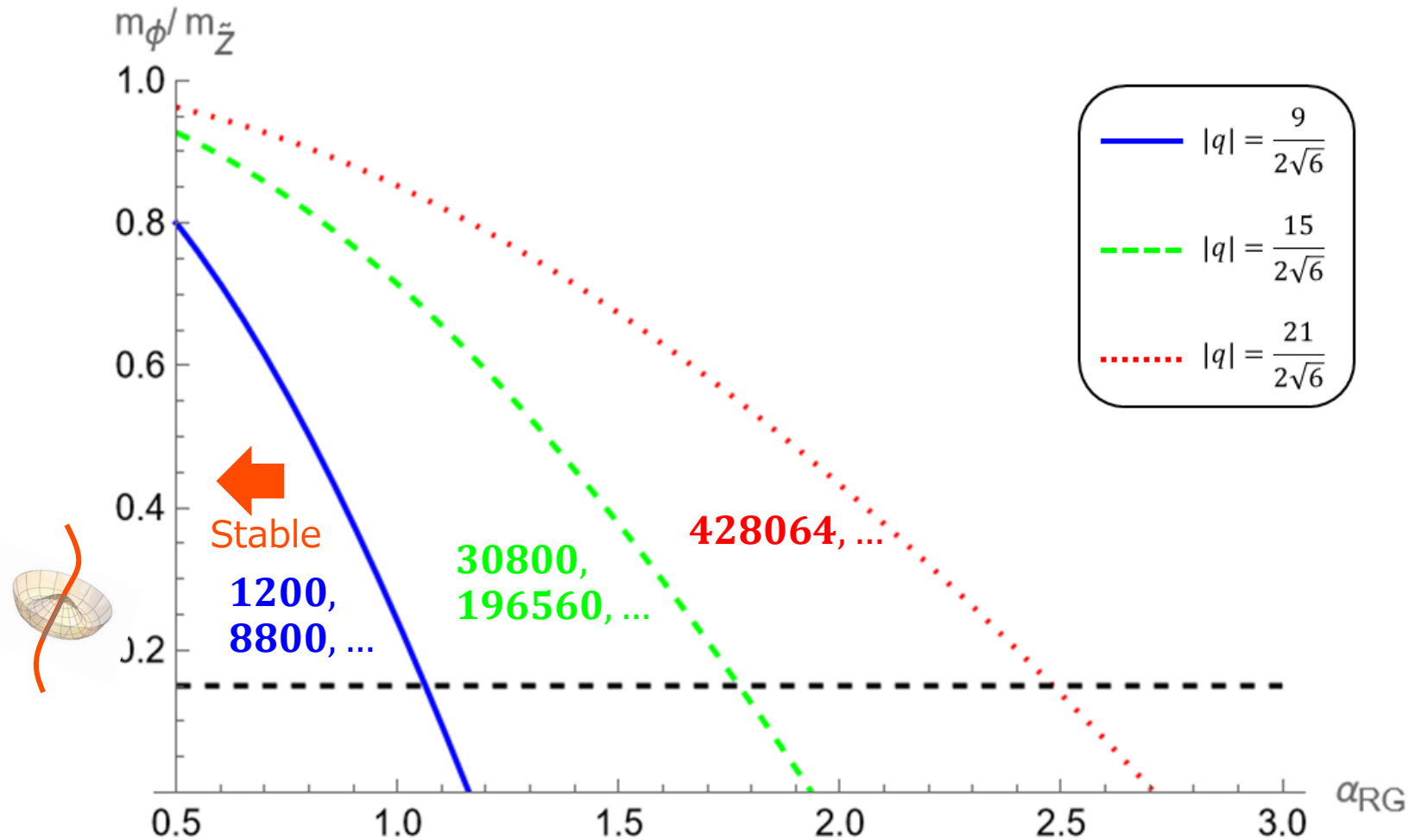
$$\mathbf{SO}(10) \rightarrow \mathbf{SU}(3)_C \times \mathbf{SU}(2)_L \times \underline{\mathbf{SU}(2)_R} \times \underline{\mathbf{U}(1)_X} \rightarrow \mathbf{SU}(3)_C \times \mathbf{SU}(2)_L \times \underline{\mathbf{U}(1)_Y}$$

\uparrow
 $\phi = (\mathbf{1}, \mathbf{1}, \mathbf{2}, q)$

$$SO(10) \rightarrow SU(3)_C \times SU(2)_L \times \underline{SU(2)_R} \times \underline{U(1)_X} \rightarrow SU(3)_C \times SU(2)_L \times \underline{U(1)_Y}$$

($\alpha_{RG} = g_{2R}/g_{1X}$ at the breaking scale)

$$\uparrow \phi = (1, 1, 2, q)$$



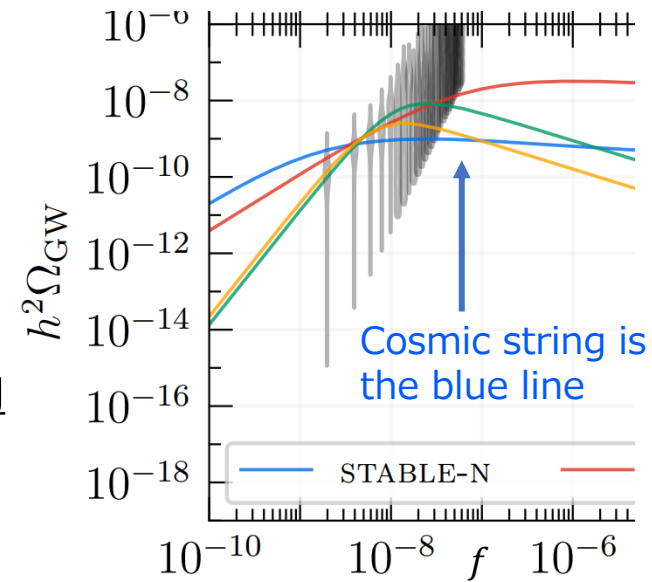
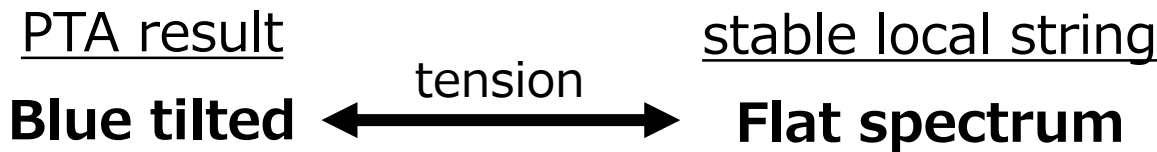
Very big representation Higgs is needed!

⇒ If generalized Z string is found, it can be a constraint for GUT.

PTA results and metastability

In June 2023, PTA collaborations reported a detection of stochastic gravitational waves.

NANOGrav [2306.16219], EPTA [2306.16214],
PPTA [2306.16215], CPTA [2306.16216]



Metastable strings can solve this tension.

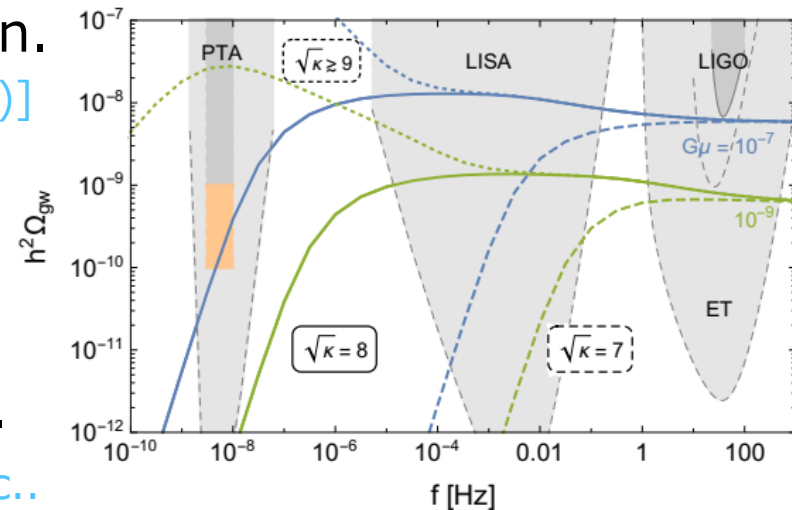
[Buchmuller, Domcke, Schmitz (2020)]

$$\text{Decay probability } P \sim \frac{\mu}{2\pi} e^{-\pi\kappa}$$

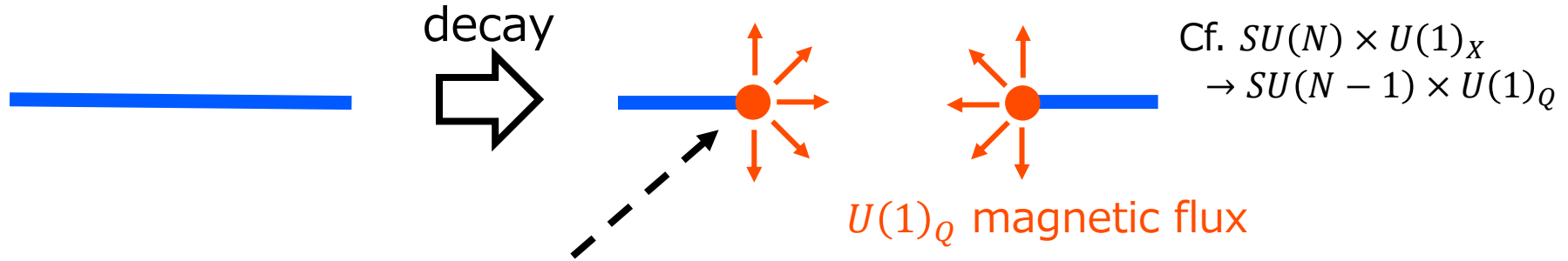
(μ : string tension)

$\sqrt{\kappa} \sim 8$ is preferred to solve the tension.

[Buchmuller, Domcke, Schmitz (2023)], etc..



Estimation of decay probability



For $N = 2$, “Nambu monopole” [Nambu (1977)]

Decay probability can be estimated as $P \sim \frac{\mu}{2\pi} e^{-\pi m^2/\mu}$ [Preskill, Vilenkin (1993)]

where Nambu-monopole mass $m \sim \frac{4\sqrt{2}\pi}{3\alpha_N} \tan^{\frac{3}{2}} \theta_G \sqrt{\frac{m_\phi}{m_{\tilde{Z}}}} v \times \mathcal{O}(1)$

String tension $\mu \sim 2\pi \frac{m_\phi}{m_{\tilde{Z}}} v^2 \times \mathcal{O}(1)$

$$\sqrt{\kappa} \sim \frac{4\sqrt{\pi}}{3\alpha_N} \tan^{\frac{3}{2}} \theta_G \sim \frac{4\sqrt{\pi}}{3\alpha_N} \left(\frac{N}{2(N-1)} \frac{m_{\tilde{Z}}^2}{m_G^2} - 1 \right)^{\frac{3}{4}} \xrightarrow{\text{e.g. } \alpha_N = 1.3, N = 7, m_G^2/m_{\tilde{Z}}^2 = 0.07} 8.1 \dots$$

Large α_N, N and $m_G/m_{\tilde{Z}}$ are preferred to explain the PTA results.

Summary

- Embedded strings are the classical solutions having 1-dimensional excited region (= cosmic string), but their stability is not guaranteed by a topological feature of the vacuum manifold.
- We have generalized the Z string to $SU(N) \times U(1)_X \rightarrow SU(N-1) \times U(1)_Q$, and studied its stability. We have found that it is needed that the $U(1)$ gauge coupling is larger than $SU(N)$ to the formation.
- We have applied the formation condition to the case that $SU(N)$ and $U(1)$ have the same origin, and found that a very large representation of scalar is needed for the formation.
- To see the relevance to the recent results of PTA collaborations, we focus on the metastability of the generalized Z string. The decay probability may be good for explaining the PTA result for large gauge couplings, mixing angle and N , but more precise calculations are needed. (Ongoing work)

Back up